

# Higher Derivative and Conformal Gravity from Bimetric and Partially Massless Bimetric Theories

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## In Collaboration with

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*arXiv:1103.6055, 1106.3344, 1109.3515,  
1109.3230, 1111.2070*
- ▶ Angnis Schmidt-May & Mikael von Strauss  
*arXiv:1203.5283, 1204.5202, 1208:1515, 1208:1797,  
1212:4525, 1303.6940*

# Outline of the talk

Review: Linear and Nonlinear massive spin-2 fields

Ghost-free bi-metric theory

Mass spectrum of bimetric theory

Partially Massless bimetric theory

Higher derivative gravity from bimetric theory

Equivalence between CG and PM bimetric theory

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# Linear massive spin-2 fields

## The Fierz-Pauli equation:

Linear massive spin-2 field,  $h_{\mu\nu}$ , in background metric  $\bar{g}_{\mu\nu}$

$$\bar{\mathcal{E}}_{\mu\nu}^{\rho\sigma} h_{\rho\sigma} - \Lambda \left( h_{\mu\nu} - \frac{1}{2} \bar{g}_{\mu\nu} h_{\rho}^{\rho} \right) + \frac{m_{\text{FP}}^2}{2} \left( h_{\mu\nu} - \bar{g}_{\mu\nu} h_{\rho}^{\rho} \right) = 0$$

[Fierz-Pauli, 1939]

- ▶ 5 propagating modes (massive spin-2)
- ▶ Massive gravity?
- ▶ What determines  $\bar{g}_{\mu\nu}$ ? (flat, dS, AdS, ...)
- ▶ Nonlinear generalizations?

[The Boulware-Deser ghost (1972)]

# Nonlinear generalizations of FP theory

- ▶ Massive gravity (fixed  $f_{\mu\nu}$ ):

$$\mathcal{L} = m_p^2 \sqrt{-g} \left[ R - m^2 V(g^{-1} f) \right]$$

- ▶ Massive spin-2 field + gravity (dynamical  $f$ ):

$$\mathcal{L} = m_p^2 \sqrt{-g} \left[ R - m^2 V(g^{-1} f) \right] + \mathcal{L}(\nabla f)(?)$$

Bimetric:  $\mathcal{L}(\nabla f) = m_f^2 \sqrt{-f} R_f(?)$

[Isham-Salam-Strathdee, 1971, 1977]

Generically, both contain a *GHOST* at the nonlinear level

[Boulware-Deser, 1972]

# Counting modes:

## Generic massive gravity:

- ▶ Linear: 5 modes
- ▶ Non-linear: 6 modes (massive spin-2 + ghost)

## Generic bimetric theory:

- ▶ Linear: 5 modes (massive,  $(\delta g - \delta f)$ )  
2 mode (massless,  $(\delta g + \delta f)$ )
- ▶ Non-linear: 7 modes + 1 (ghost)

Complication: Since the ghost shows up nonlinearly, its absence needs to be established nonlinearly

# Construction of ghost-free nonlinear theories

**Based on “Decoupling limit” (perturbative):**

A specific  $V(\sqrt{g^{-1}}\eta)$  was obtained and shown to be ghost-free in a “decoupling limit” and also perturbatively in  $h = g - \eta$

*[de Rham, Gabadadze, 2010; de Rham, Gabadadze, Tolley, 2010]*

**Non-linear Hamiltonian methods (non-perturbative):**



# Construction of ghost-free nonlinear theories

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## Non-linear Hamiltonian methods (non-perturbative):

Questions not answerable by “decoupling limit”:

- ▶ Is massive gravity with  $V(\sqrt{g^{-1}\eta})$  ghost-free nonlinearly?

*[SFH, Rosen (1106.3344, 1111.2070)]*

- ▶ Is it ghost-free for generic fixed  $f_{\mu\nu}$ ?

*[SFH, Rosen, Schmidt-May (1109.3230)]*

(see Cedric’s talk for alternative approaches)

- ▶ Can  $f_{\mu\nu}$  be given ghost-free dynamics?

*[SFH, Rosen (1109.3515)]*

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## Ghost-free bimetric theory

**Digression:** Elementary symmetric polynomials of  $\mathbb{X}$  with eigenvalues  $\lambda_1, \dots, \lambda_4$ :

$$\begin{aligned}e_0(\mathbb{X}) &= 1, & e_1(\mathbb{X}) &= \lambda_1 + \lambda_2 + \lambda_3 + \lambda_4, \\e_2(\mathbb{X}) &= \lambda_1\lambda_2 + \lambda_1\lambda_3 + \lambda_1\lambda_4 + \lambda_2\lambda_3 + \lambda_2\lambda_4 + \lambda_3\lambda_4, \\e_3(\mathbb{X}) &= \lambda_1\lambda_2\lambda_3 + \lambda_1\lambda_2\lambda_4 + \lambda_1\lambda_3\lambda_4 + \lambda_2\lambda_3\lambda_4, \\e_4(\mathbb{X}) &= \lambda_1\lambda_2\lambda_3\lambda_4 = \det \mathbb{X}.\end{aligned}$$

$$\begin{aligned}e_0(\mathbb{X}) &= 1, & e_1(\mathbb{X}) &= [\mathbb{X}], \\e_2(\mathbb{X}) &= \frac{1}{2}([\mathbb{X}]^2 - [\mathbb{X}^2]), \\e_3(\mathbb{X}) &= \frac{1}{6}([\mathbb{X}]^3 - 3[\mathbb{X}][\mathbb{X}^2] + 2[\mathbb{X}^3]), \\e_4(\mathbb{X}) &= \frac{1}{24}([\mathbb{X}]^4 - 6[\mathbb{X}]^2[\mathbb{X}^2] + 3[\mathbb{X}^2]^2 + 8[\mathbb{X}][\mathbb{X}^3] - 6[\mathbb{X}^4]), \\e_k(\mathbb{X}) &= 0 \quad \text{for } k > 4,\end{aligned}$$

$$[\mathbb{X}] = \text{Tr}(\mathbb{X}), \quad e_n(\mathbb{X}) \sim (\mathbb{X})^n$$

- ▶ The  $e_n(\mathbb{X})$ 's and  $\det(\mathbb{1} + \mathbb{X})$ :

$$\begin{aligned} \det(\mathbb{1} + \mathbb{X}) &= \sum_{n=0}^4 e_n(\mathbb{X}) \\ &= \sum_{n=0}^4 \frac{-1}{n!(4-n)!} \epsilon_{\mu_1 \dots \mu_n \lambda_{n+1} \dots \lambda_4} \epsilon^{\nu_1 \dots \nu_n \lambda_{n+1} \dots \lambda_4} \mathbb{X}_{\nu_1}^{\mu_1} \dots \mathbb{X}_{\nu_n}^{\mu_n} \end{aligned}$$

- ▶ Introduce “deformed determinant” :

$$\widehat{\det}(\mathbb{1} + \mathbb{X}) = \sum_{n=0}^4 \beta_n e_n(\mathbb{X})$$

- ▶ Observation:

$$V(\sqrt{g^{-1}f}) = \widehat{\det}(\mathbb{1} + \sqrt{g^{-1}f})$$

[SFH & R. A. Rosen (1103.6055)]

# Ghost-free bi-metric theory

Ghost-free combination of kinetic and potential terms for  $g$  &  $f$ :

$$\mathcal{L} = m_g^2 \sqrt{-g} R_g - 2m^4 \sqrt{-g} \sum_{n=0}^4 \beta_n e_n(\sqrt{g^{-1}f}) + m_f^2 \sqrt{-f} R_f$$

Note,

$$\sqrt{-g} \sum_{n=0}^4 \beta_n e_n(\sqrt{g^{-1}f}) = \sqrt{-f} \sum_{n=0}^4 \beta_{4-n} e_n(\sqrt{f^{-1}g})$$

Hamiltonian analysis: **7** nolinear propagating modes, **no ghost!**

$$C(\gamma, \pi) = 0, \quad C_2(\gamma, \pi) = \frac{d}{dt} C(x) = \{H, C\} = 0$$

[SFH, Rosen (1109.3515, 1111.2070)]

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# Mass spectrum of bimetric theory

[SFH, A. Schmidt-May, M. von Strauss 1208:1515, 1212:4525]

$$S_{gf} = - \int d^d x \left[ m_g^{d-2} \sqrt{g} R_g - 2m^d \sqrt{g} \sum_{n=0}^d \beta_n e_n(S) + m_f^{d-2} \sqrt{f} R_f \right]$$

Equations of motion:

$$R_{\mu\nu}(g) - \frac{1}{2} g_{\mu\nu} R(g) + V_{\mu\nu}^g = T_{\mu\nu}^g$$

$$R_{\mu\nu}(f) - \frac{1}{2} f_{\mu\nu} R(f) + V_{\mu\nu}^f = T_{\mu\nu}^f$$

(for classical solutions, see talk by Mikhail Volkov)

- ▶ When are the 7 fluctuations in  $\delta g_{\mu\nu}$ ,  $\delta f_{\mu\nu}$  good mass eigenstates? (well-defined FP mass)
- ▶ How to characterize deviations from General Relativity?

# Mass spectrum: proportional backgrounds

FP masses exist only around,

$$\bar{f}_{\mu\nu} = c^2 \bar{g}_{\mu\nu}$$

$g$  and  $f$  equations:

$$R_{\mu\nu}(\bar{g}) - \frac{1}{2} \bar{g}_{\mu\nu} R(\bar{g}) + \begin{pmatrix} \Lambda_g \\ \Lambda_f \end{pmatrix} \bar{g}_{\mu\nu} = 0 \text{ or } \begin{pmatrix} T_{\mu\nu}^g \\ T_{\mu\nu}^f \end{pmatrix}$$

$$\Lambda_g = \frac{m^d}{m_g^{d-2}} \sum_{k=0}^{d-1} \binom{d-1}{k} c^k \beta_k, \quad \Lambda_f = \frac{m^d}{m_f^{d-2}} \sum_{k=1}^d \binom{d-1}{k-1} c^{k+2-d} \beta_k$$

**Implication:**

$$\Lambda_g = \Lambda_f \quad \Rightarrow \quad c = c(\beta_n, \alpha \equiv m_f/m_g)$$

(Exception: Partially massless (PM) theory)



# Mass spectrum around proportional backgrounds

Linear modes:

$$\delta M_{\mu\nu} = \frac{1}{2c} \left( \delta f_{\mu\nu} - c^2 \delta g_{\mu\nu} \right), \quad \delta G_{\mu\nu} = \left( \delta g_{\mu\nu} + \alpha^{d-2} c^{d-4} \delta f_{\mu\nu} \right)$$

$$\bar{\mathcal{E}}_{\mu\nu}^{\rho\sigma} \delta G_{\rho\sigma} - \Lambda_g \left( \delta G_{\mu\nu} - \frac{1}{2} \bar{g}_{\mu\nu} \bar{g}^{\rho\sigma} \delta G_{\rho\sigma} \right) = 0,$$

$$\begin{aligned} \bar{\mathcal{E}}_{\mu\nu}^{\rho\sigma} \delta M_{\rho\sigma} - \Lambda_g \left( \delta M_{\mu\nu} - \frac{1}{2} \bar{g}_{\mu\nu} \bar{g}^{\rho\sigma} \delta M_{\rho\sigma} \right) \\ + \frac{1}{2} m_{\text{FP}}^2 \left( \delta M_{\mu\nu} - \bar{g}_{\mu\nu} \bar{g}^{\rho\sigma} \delta M_{\rho\sigma} \right) = 0 \end{aligned}$$

The FP mass of  $\delta M$ :

$$m_{\text{FP}}^2 = \frac{m^d}{m_g^{d-2}} \left( 1 + (\alpha c)^{2-d} \right) \sum_{k=1}^{d-1} \binom{d-2}{k-1} c^k \beta_k$$

# Bimetric as massive spin-2 field + gravity

Nonlinear extensions of linear modes:

$$G_{\mu\nu} = g_{\mu\nu} + c^{d-4} \alpha^{d-2} f_{\mu\nu}, \quad M_{\mu\nu}^G = G_{\mu\rho} (\sqrt{g^{-1}f})^\rho{}_\nu - c G_{\mu\nu}$$

$G$ : No ghost-free matter coupling, not the gravitational metric.

- ▶ Bimetric as gravity + massive spin-2 field:

$$g_{\mu\nu}, \quad M_{\mu\nu} = g_{\mu\rho} (\sqrt{g^{-1}f})^\rho{}_\nu - c g_{\mu\nu}, \quad m_g \gg m_f$$

(see talk by Keisuke Izumi)

- ▶  $M = 0 \Rightarrow$  GR.
- ▶  $M \neq 0 \Rightarrow$  deviations from GR, driven by matter couplings.

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# Partial masslessness in FP theory

$$\bar{\mathcal{E}}^{\rho\sigma}_{\mu\nu} h_{\rho\sigma} - \Lambda(h_{\mu\nu} - \frac{1}{2}\bar{g}_{\mu\nu}h^\rho{}_\rho) + \frac{m_{FP}^2}{2}(h_{\mu\nu} - \bar{g}_{\mu\nu}h^\rho{}_\rho) = 0$$

dS/Einstein backgrounds:

$$\bar{g}_{\mu\nu} : \quad R_{\mu\nu} - \frac{1}{2}g_{\mu\nu}R + \Lambda g_{\mu\nu} = 0$$

Higuchi Bound:

$$m_{FP}^2 = \frac{2}{3}\Lambda$$

New gauge symmetry:

$$\Delta h_{\mu\nu} = (\nabla_\mu \nabla_\nu + \frac{\Lambda}{3})\xi(x)$$

Gives  $5-1=4$  propagating modes

[Deser, Waldron, ... (1983-2012)]

Can a nonlinear extension of PM theory exist?

# Partial masslessness beyond FP theory

Non-linear PM theory = Nonlinear spin-2 fields with a gauge invariance!

Does it exist? Independent of dS/Einstein backgrounds?

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## Known perturbative results around dS:

- ▶ Cubic PM vertices ( $\sim h^3$ ) in  $d = 4$  *[Zinoviev (2006)]*
- ▶ Cubic PM vertices exist only in  $d = 3, 4$  with 2 derivatives  
For  $d > 4$ , higher derivative theory needed.  
*[Joung, Lopez, Taronna (2012)]*

We will identify a specific bimetric theory as the candidate nonlinear PM theory

# Partial masslessness in Bimetric theory

[SFH, Schmidt-May, von Strauss, 1208:1797, 1212:4525]

- 1) Assume a nonlinear bimetric theory with PM symmetry exists
- 2) Around  $\bar{f} = c^2 \bar{g}$ ,  $\delta M_{\mu\nu} \sim h_{\mu\nu}$  satisfies the FP equation. Then the action of symmetry on  $\delta M_{\mu\nu}$  &  $\delta G_{\mu\nu}$  must be:

$$\delta M_{\mu\nu} \rightarrow \delta M_{\mu\nu} + \left( \nabla_\mu \nabla_\nu + \frac{\Lambda}{3} \bar{g}_{\mu\nu} \right) \xi(x), \quad \delta G_{\mu\nu} \rightarrow \delta G_{\mu\nu}$$

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- ▶ Find the transformation of  $\delta g_{\mu\nu}$  &  $\delta f_{\mu\nu}$ .
- ▶ Shift the transf. to dynamical backgrounds  $\bar{g}_{\mu\nu}$  &  $\bar{f}_{\mu\nu}$
- ▶ For the dS-preserving subset  $\xi = \xi_0$  (cont), this gives,

$$\bar{g}'_{\mu\nu} = (1 + a\xi_0) \bar{g}_{\mu\nu}, \quad \bar{f}'_{\mu\nu} = (1 + b\xi_0) \bar{f}_{\mu\nu}$$

$$\bar{f}' = c'^2(\xi_0) \bar{g}' \quad c' \neq c$$

A symmetry can exist only if  $\Lambda_g = \Lambda_f$  does not determine  $c$



# Candidate PM bimetric theory in d=4

The necessary condition for the existence of PM symmetry is that  $c$  is **not** determined by  $\Lambda_g = \Lambda_f$ , or

$$\beta_1 + \left(3\beta_2 - \alpha^2\beta_0\right) c + \left(3\beta_3 - 3\alpha^2\beta_1\right) c^2 + \left(\beta_4 - 3\alpha^2\beta_2\right) c^3 + \alpha^2\beta_3c^4 = 0$$

This gives the candidate nonlinear PM theory (d=4)

$$\alpha^2\beta_0 = 3\beta_2, \quad 3\alpha^2\beta_2 = \beta_4, \quad \beta_1 = \beta_3 = 0$$

# Nonlinear PM bimetric theory

## Checks:

- ▶  $m_{\text{FP}}^2 = 2 \frac{m^4}{m_g^2} (\alpha^{-2} + c^2) \beta_2 = \frac{2}{3} \Lambda_g$
- ▶ For  $d > 4$ , all  $\beta_n = 0$ . Nonlinear PM bimetric exists only for  $d = 3, 4$ .
- ▶ In  $d > 4$  PM is restored by Lanczos-Lovelock terms
- ▶ Realization of the  $\xi_0$  gauge transformation in the nonlinear theory.

Full Gauge symmetry of the nonlinear theory? (not yet known, but ....)

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# Higher derivative gravity and Conformal gravity

HD gravity:

$$S_{(2)}^{\text{HD}}[g] = m_g^2 \int d^4x \sqrt{g} \left[ \Lambda + c_R R(g) - \frac{c_{RR}}{m^2} \left( R^{\mu\nu} R_{\mu\nu} - \frac{1}{3} R^2 \right) \right]$$

7 modes: 2 (massless spin-2) + 5 (massive spin-2 **ghost**)  
[Stelle (1977)]

Conformal Gravity:

$$S^{\text{CG}}[g] = -c \int d^4x \sqrt{g} \left[ R^{\mu\nu} R_{\mu\nu} - \frac{1}{3} R^2 \right]$$

Invariant under Weyl scalings  $\Rightarrow$

6 modes: 2 (massless spin-2) + 4 **ghost** modes

[Riegert (1984), Maldacena (2011)]

# Conformal gravity and PM theory

- ▶ Maldacena: CG spectrum in dS background  $\sim$  linear PM. CG nonlinear PM candidate?
- ▶ Deser-Waldron: No PM spectrum away from dS backgrounds (caveate: too restrictive condition on the spectrum)

# HD gravity from Bimetric theory

Define

$$S = \sqrt{g^{-1}f}, \quad P_{\mu\nu} = R_{\mu\nu} - \frac{1}{2(d-1)}g_{\mu\nu}R$$

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$$S = \sqrt{g^{-1}f}, \quad P_{\mu\nu} = R_{\mu\nu} - \frac{1}{2(d-1)}g_{\mu\nu}R$$

Solve the bimetric  $g_{\mu\nu}$  equation algebraically for  $f_{\mu\nu}$ , as an expansion in  $R_{\mu\nu}(g)/m^2$ ,

$$S^\mu{}_\nu = a\delta^\mu{}_\nu + \frac{a_1}{m^2}P^\mu{}_\nu + \frac{a_2}{m^4} \left[ \left( P^\mu{}_\nu{}^2 - P P^\mu{}_\nu \right) + \frac{1}{d-1}e_2(P)\delta^\mu{}_\nu \right] + \mathcal{O}(m^{-6})$$

Compute  $f = f(g)$ . Then

$$S^{\text{BM}}[g, f(g)] = S^{\text{HD}}[g]$$

# HD gravity from Bimetric theory

- ▶ 4-derivative ( $\sim R^2$ ) truncation:

$$S_{(2)}^{\text{BM}}[g, f(g)] = S_{(2)}^{\text{HD}}[g]$$

The spin-2 ghost in 4-derivative HD gravity is an artifact of this truncation (can be illustrated in a linear theory).

- ▶ The correspondence is not an equivalence of the truncated theories (in general). Different truncated EoM's.
- ▶ For PM bimetric theory one obtains conformal gravity.



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- ▶ In PM bimetric theory, determine  $f_{\mu\nu}$  from  $g$ -equation. Substitute in  $f$ -equation (not in the action) to get,

$$B_{\mu\nu} + \mathcal{O}(R^3/m^2) = 0$$

In the low curvature limit, PM bimetric theory has a gauge symmetry even away from dS and definitely propagates  $7 - 1 = 6$  modes! None is a ghost

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- ▶ CG eom is the low curvature limit of PM bimetric eom.  
**Conversely, PM bimetric is a ghost-free completion of CG**